

# Statistical Analysis of Different $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ Searches

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## Abstract

A combined statistical analysis of the experimental results of the LSND and KARMEN  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  oscillation search is presented. LSND has evidence for neutrino oscillations that is not confirmed by the KARMEN experiment. This joint analysis is based on the final likelihood results for both data sets. A frequentist approach is applied to deduce confidence regions. At a combined confidence level of 36 %, there is no area of oscillation parameters compatible with both experiments. For the complementary confidence of  $1 - 0.36 = 64$  %, there are two well defined regions of oscillation parameters  $(\sin^2(2\Theta), \Delta m^2)$  compatible with both experiments.

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## I. INTRODUCTION

Currently, there are three neutrino anomalies interpreted as evidence for neutrino oscillations, namely the atmospheric [1], the solar [2] and the LSND [3] anomaly with three distinct squared neutrino mass differences. However, oscillations between the three Standard Model (SM) neutrinos are described by only two independent neutrino mass-squared differences, allowing only two of the above anomalies as being due to oscillations in a SM scenario extended by massive neutrinos. Before enlarging the neutrino sector by some additional, unknown *sterile* neutrinos, each of the experimental results has to be checked independently.

Over the last years, the controversial results of the two experiments LSND (Liquid Scintillator Neutrino Detector at LANSCE, Los Alamos, USA) and KARMEN (KARlsruhe Rutherford Medium Energy Neutrino experiment at ISIS, Rutherford, UK) both searching for neutrino oscillations  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  in a *short baseline* regime, have led to intense discussions. The two experiments are similar as they use  $\bar{\nu}_\mu$  beams from the  $\pi^+ \rightarrow \mu^+ + \nu_\mu$  decay at rest (DAR) chain  $\pi^+ \rightarrow \mu^+ + \nu_\mu$  followed by  $\mu^+ \rightarrow e^+ + \nu_e + \bar{\nu}_\mu$  with energies up to 52 MeV. Furthermore, both experiments are looking for  $\bar{\nu}_e$  from  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  oscillations according to the two-flavor oscillation probability [29]

$$P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) = \sin^2(2\Theta) \sin^2 \left( 1.27 \frac{\Delta m^2 \cdot L}{E_\nu} \right) \quad (1.1)$$

with the difference of the squared masses  $\Delta m^2$  in  $\text{eV}^2/c^4$ , the flight length of the neutrino  $L$  in meters and the neutrino energy  $E_\nu$  in units of MeV. The detection reaction  $p(\bar{\nu}_e, e^+)n$  provides a spatially correlated delayed coincidence signature of a prompt  $e^+$  and a subsequent neutron capture signal.

This paper describes a combined statistical analysis of the final LSND and KARMEN 2 results. There are significant changes to the treatment of LSND data in the final LSND paper [3]. The present paper uses only the decay-at-rest (DAR)  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  data, and leaves out the impact of the decay-in-flight (DIF)  $\nu_\mu \rightarrow \nu_e$  data. As shown later, the LSND data has a favored region in the  $7 \text{ eV}^2/c^4$  region of  $\Delta m^2$  as a consequence of this change. This is in contrast to the LSND result [3] which disfavors this solution at 90% C.L. because of the inclusion of the DIF flux. A motivation for this change is that a direct comparison of the two experiments is less model dependent if one considers the LSND DAR flux only, since KARMEN does not have a DIF neutrino flux present in its source.

The analysis is based on a frequentist approach following the suggestions of [4]. For both experiments, the data are analysed with a maximum likelihood analysis followed by the extraction of confidence levels in a unified approach. It is not the task nor the purpose of this analysis to treat the apparent disagreement between the two experiments. Rather, assuming no serious systematical error in either experiment or their interpretation with respect to  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  oscillation, we apply a consistent statistical analysis to quantitatively establish their level of compatibility. The work follows an earlier analysis [5] based on intermediate data sets.

There are two separate questions investigated in this paper that relate to the LSND and KARMEN data. The first question is: at what level are the two data sets *compatible* with each other? This question addresses whether or not the two experiments force one to draw opposite conclusions from their respective data.

If one *assumes* that the data sets are compatible, the second question becomes: what are the most likely regions for the oscillation parameters? This method is somewhat analogous to treating the KARMEN data as coming from a 'near detector' at 17 meters, and treating the LSND data as coming from a 'far detector' at 30 meters. It ignores the question of systematic differences between the experiments, such as whether or not, for example, an unaccounted source of  $\bar{\nu}_e$  background neutrinos was present in the LSND data and not present in the KARMEN data.

The paper is organized as follows: After a short overview of the experimental setups of both experiments, section II describes the data analyses leading to likelihood functions of the oscillation parameters. Section III is devoted to the extraction of confidence regions for the individual experiments. In section IV, the experimental results are combined statistically, extracting levels of compatibility as well as oscillation parameters compatible with both experiments. We conclude with implications of this joint analysis, a comparison with other oscillation searches and an outlook to upcoming experimental tests of the favored oscillation parameters presented in this analysis.

## A. The LSND experiment

The source of neutrinos for the LSND experiment was the interaction of the 798 MeV proton beam at the Los Alamos Neutron Science Center (LANSCE), in which a large number

of pions, mostly  $\pi^+$  are produced. The  $\pi^-$  are mainly absorbed and only a small fraction decay to  $\mu^-$ , which in turn are largely captured. Thus, the resulting neutrino source is dominantly due to  $\pi^+ \rightarrow \mu^+ + \nu_\mu$  and  $\mu^+ \rightarrow e^+ + \nu_e + \bar{\nu}_\mu$  decays, most of which decay at rest (DAR). With a very small contamination of  $\bar{\nu}_e/\bar{\nu}_\mu \sim 8 \cdot 10^{-4}$ , a measurement of the reaction  $\bar{\nu}_e + p \rightarrow n + e^+$  provides a sensitive way to search for  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  oscillations. Such events are identified by detection of both the  $e^+$  and the 2.2 MeV  $\gamma$  from the reaction  $p(n, \gamma)d$ . In addition, the  $\nu_e$  flux from  $\pi^+$  and  $\mu^+$  decay-in-flight (DIF) is very small, which allows a search for  $\nu_\mu \rightarrow \nu_e$  oscillations via the measurement of electrons above the Michel electron endpoint from the reaction  $^{12}\text{C}(\nu_e, e^-)^{12}\text{N}$ .

The LSND experiment took data over six calendar years (1993-1998). During this period the LANSCE accelerator operated for 17 months, delivering 28 896 C of protons on the production target. The LSND detector [6] consisted of an approximately cylindrical tank 8.3 m long by 5.7 m in diameter. The center of the detector was located 30 m from the beam stop neutrino source. The tank was filled with liquid scintillator consisting of mineral oil and 0.031 g/l of b-PBD. This low scintillator concentration allowed the detection of both Čerenkov light and scintillation light. Photomultiplier time and pulse-height signals were used to reconstruct the track with an average RMS position resolution of  $\sim 14$  cm and an energy resolution of  $\sim 7\%$  at the Michel endpoint of 52.8 MeV.

## B. The KARMEN experiment

The KARMEN experiment used as neutrino source the pulsed spallation neutron source ISIS of the Rutherford Appleton Laboratory delivering 800 MeV protons. In contrast to the LANSCE source, the protons are extracted from the synchrotron as an intense but narrow double pulse, consisting of two parabolic pulses with a width of 100 ns separated by a peak-to-peak gap of 325 ns. The unique time structure of the ISIS proton pulses allowed a clear separation of neutrino-induced events from any beam unrelated background. The intrinsic contamination of the ISIS  $\bar{\nu}_\mu$ -beam is  $\bar{\nu}_e/\bar{\nu}_\mu = 6.4 \cdot 10^{-4}$  [7].

The KARMEN detector [8] was a rectangular high resolution liquid scintillation calorimeter, located at a mean distance of 17.7 m from the ISIS target at an angle of 100 degrees relative to the proton beam. The liquid scintillator was enclosed by a multilayer active veto system and a 7000 t steel shielding. The 65 m<sup>3</sup> of liquid scintillator consisted of a

mixture of paraffin oil (75%vol.), Pseudocumene (25%vol.) and 2 g/l scintillating additive 1-phenyl-3-mesityl-2-pyrazoline. The liquid scintillator volume was optically separated into 512 independent central modules. Gadolinium was implemented between the module walls for an efficient detection of thermal neutrons  $\text{Gd}(n,\gamma)$  with on average 3  $\gamma$ 's of energy  $\sum E_\gamma = 8 \text{ MeV}$ . The KARMEN detector as liquid scintillator calorimeter was optimized for high energy resolution of  $\sigma_E = 11.5\%/\sqrt{E(\text{MeV})}$ . Each event had energy, time and position information, as well as the number of addressed modules and their relative time differences.

The KARMEN 2 experiment took data from February 1997 to March 2001. During this time, protons equivalent to a total charge of 9425 Coulombs have been accumulated on the ISIS target. This corresponds to  $2.71 \cdot 10^{21} \bar{\nu}_\mu$  from the ISIS beam stop target.

## II. DATA ANALYSIS AND RESULTS AS LIKELIHOOD FUNCTIONS

For both experiments, a  $\bar{\nu}_e$  signal from  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  oscillations consists of a spatially correlated delayed  $(e^+, n)$  sequence from  $p(\bar{\nu}_e, e^+)n$ . Due to the different experimental techniques, the parameters to identify such sequences vary. Details of the LSND event reconstruction can be found in [9] and [6], the requirements for event sequences in KARMEN are described in [10].

In LSND, a  $(e^+, n)$  sequence requires a prompt 'electron-like' event with energy  $E_e > 20 \text{ MeV}$  followed by a low energy  $\gamma$ -event. The information about the delayed event is characterized via a likelihood ratio  $R_\gamma$ : If within one millisecond after the initial event another event is recorded at distance  $\Delta r \leq 250 \text{ cm}$ , the likelihood ratio  $R_\gamma$  in energy (PMT hits), time and distance of being a correlated  $p(n, \gamma)$  over an accidental coincidence is calculated, otherwise  $R_\gamma = 0$ . Requiring a high likelihood ratio  $R_\gamma$  selects  $(e^+, n)$  correlated events with low uncorrelated background, the so-called "gold-plated" events. However, to determine the oscillation parameters  $\sin^2(2\Theta)$  and  $\Delta m^2$  in a likelihood analysis, an event sample with a much looser cut in  $R_\gamma$  is used, leading to higher efficiency for the oscillation channel, but naturally increasing the background. For the LSND sample analysed here, all the cuts described in [3] have been applied with two exceptions: The energy of the prompt event has been restricted to  $20 < E_e < 60 \text{ MeV}$  instead of  $20 < E_e < 200 \text{ MeV}$  and we applied a loose cut of  $R_\gamma > 10^{-5}$  instead of no  $R_\gamma$ -cut at all. The reason for these changes is twofold. First, the upper energy cut  $E_e < 60 \text{ MeV}$  excludes any significant influence of the

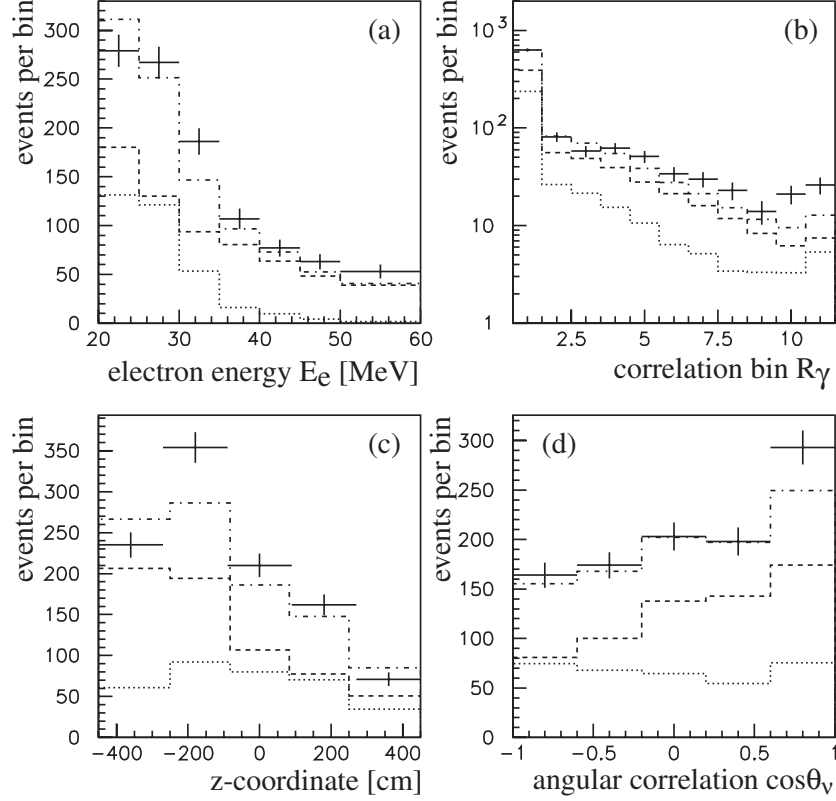


FIG. 1: LSND data ensemble, together with the BUB (dashed), BRB (dotted) and total background expectations BUB+BRB (dashed-dotted): (a) electron energy, (b)  $R_\gamma$ , (c) spatial distribution along detector axis, and (d) directional angle  $\cos \theta_\nu$ .

$\nu_\mu \rightarrow \nu_e$  channel on the much more sensitive  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  channel. For any value of the oscillation parameter  $\Delta m^2$ , the expected contribution from  $\nu_\mu \rightarrow \nu_e$  remains below 2% of the  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  signal. The second reason is a more technical one. Applying a unified frequentist approach to the likelihood function gained from the reduced data implies simulation and analysis of a large number of event samples analog to the experiment. Therefore, one is interested not to have too large samples containing mainly background events. With the cuts above, the original event sample is reduced by a factor of 5.5 to 1032 candidates, with an efficiency for  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  at high  $\Delta m^2$  reduced to 64.4% of the original one (only 3.2% of the expected  $\nu_\mu \rightarrow \nu_e$  events remain).

Figure 1 shows the event sample comprising 1032 beam-on events. Four variables are used to categorize the events: The energy of the primary electron  $E_e$ , its spatial distribution along the detector axis  $z$  [30] and the angle  $\cos \theta_\nu$  between the direction of the incident

TABLE I: Contributions to the LSND data sample of different beam-related (BRB 3–6,8) and beam-unrelated (BUB) background processes as well as the expected oscillation events for  $\sin^2(2\Theta) = 1$ ,  $\Delta m^2 = 100 \text{ eV}^2/c^4$ .

Contribution	Signal or Background Source	Process	Expected Number of Events
1	$\bar{\nu}_\mu \rightarrow \bar{\nu}_e$	$\bar{\nu}_e p \rightarrow e + n$	11350±115
2	BUB		635.0±26.2
3	DAR $\nu_e$	$\nu_e {}^{12}\text{C} \rightarrow e^- N_{g.s.}$ $\nu_e {}^{12}\text{C} \rightarrow e^- N^*$ $\nu_e {}^{13}\text{C} \rightarrow e^- N$ $\nu e \rightarrow \nu e$	312.2±18.5
4	DIF $\nu_\mu$	$\nu_\mu C \rightarrow \mu^- N^*$ $\nu_\mu C \rightarrow \mu^- N_{g.s.}$	7.4
5	DIF $\bar{\nu}_\mu$	$\bar{\nu}_\mu p \rightarrow \mu^+ n$ $\bar{\nu}_\mu C \rightarrow \mu^+ B^*$ $\bar{\nu}_\mu C \rightarrow \mu^+ B_{g.s.}$	3.9
6	DAR $\bar{\nu}_e$ ( $\mu^-$ DAR)	$\bar{\nu}_e p \rightarrow e^+ n$	12.4
7	$\nu_\mu \rightarrow \nu_e$	$\nu_e C \rightarrow e^- N$	227±30
8	DIF $\pi^+ \rightarrow \nu_e$ and $\mu^+ \rightarrow \nu_e$ decay	$\nu_e C \rightarrow e^- N$	1.0

neutrino and the reconstructed electron path. The fourth variable is the likelihood ratio  $R_\gamma$  for a ( $e^+, n$ ) coincidence. The binning of this distribution with the upper edge of bin number  $N$  representing a value  $R_\gamma = 10^{N/3-1}$  has been chosen for practical reasons. Note that the spectra in Figure 1 are projections of a 4-dim space of correlated parameters for each event. Superimposed to the data are shown the beam-unrelated and beam-related background contributions amounting to a total number of  $N_{bgd}^{tot} = 971.9$  events. The expected background contributions are broken down in Table I, for further details we refer to [3].

Formally, each beam-on event  $j$  of the  $N_{beam-on} = 1032$  candidates is assigned a probability  $p_j(\vec{x})$  equal to a sum of probabilities  $q_i(\vec{x})$  from the backgrounds plus oscillations. The vector  $\vec{x}$  describes herein the spectral parameters of each event  $j$ ,  $\vec{x}_j = (E_{ej}, R_{\gamma j}, \cos \theta_{\nu j}, z_j)$ . It then remains to add the  $q_i$  with expected fractional contributions  $r_i$  and take the product

over all the beam-on events. The likelihood is thus

$$\mathcal{L} = \left( \prod_{j=1}^{N_{beam-on}} p_j \right), \quad (2.1)$$

where

$$p_j(\vec{x}_j) = \sum_{i=1}^{N_{contr.}} q_i(\vec{x}_j) \cdot r_i. \quad (2.2)$$

Additionally, two normalization requirements must hold:

$$\sum_{i=1}^{N_{contr.}} r_i = 1, \quad (2.3)$$

and

$$\int dE_e dR_\gamma d(\cos \theta_\nu) dz q_i(E_e, R_\gamma, \cos \theta_\nu, z) = 1 \quad (2.4)$$

for each contribution,  $i$ . Together, these requirements ensure that every observed beam-on event has a probability of occurrence equal to 1.

To allow for the fact that the backgrounds are known to some limited accuracy only, the background is varied by calculating the above likelihood at each point in the  $(\sin^2(2\Theta), \Delta m^2)$  plane many times, varying over the expected  $\sigma$  for each background. For each background configuration, the  $\mathcal{L}$  is weighted with a Gaussian factor for each background that is off its central value. The background variations are performed in a simplified manner, i.e. the beam-unrelated background (BUB) varies independently and all the beam-related backgrounds (BRBs) are locked together. Finally, the likelihood can be expressed as

$$\mathcal{L} = \int \mathcal{D}N_{bgd} \exp(-(N_{bgd} - N_{bgd,exp})^2/2\sigma^2) \cdot \prod_{j=1}^{N_{beam-on}} p_j, \quad (2.5)$$

where the  $\int \mathcal{D}N_{bgd}$  represents, schematically, the background variation described above. Integrating over different background contributions eventually reduces the free parameters of the likelihood procedure from  $N_{contr.} + 1$  to the 2 free parameters of real interest within this analysis, the oscillation parameters:

$$\mathcal{L}(r_1, \dots, r_{N_{contr.}-1}, \sin^2(2\Theta), \Delta m^2) \rightarrow \mathcal{L}(\sin^2(2\Theta), \Delta m^2) \quad (2.6)$$

Figure 2 shows the logarithm of the event-based likelihood function,  $\ln \mathcal{L}(\sin^2(2\Theta), \Delta m^2)$ , for the 1032 events investigated. The maximum is reached at a parameter combination  $(\sin^2(2\Theta)=0.85, \Delta m^2=0.055 \text{ eV}^2/\text{c}^4)$  corresponding to a total oscillation signal of  $N_{osc} =$



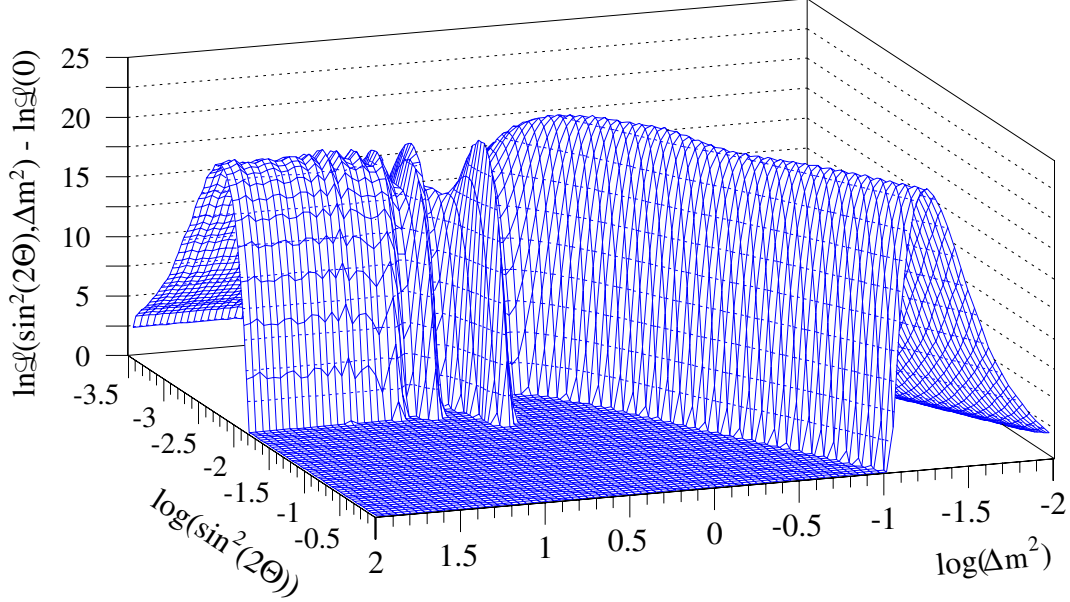


FIG. 2: LSND logarithmic likelihood function  $\ln \mathcal{L}(\sin^2(2\Theta), \Delta m^2)$

$N(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) + N(\nu_\mu \rightarrow \nu_e) = 66.4 + 0.3 = 66.7$ . The difference to the no-oscillation hypothesis as expressed in logarithmic likelihood units is

$$\ln \mathcal{L}(0.85, 0.055 \text{eV}^2/c^4) - \ln \mathcal{L}(0, 0) = 23.5 \quad (2.7)$$

underlining the significance of the additional  $\bar{\nu}_e$  signal among the event sample. Note that the extracted oscillation events  $N_{osc}$  scale with the efficiency of the applied cuts and are in good agreement with the best fit result stated in [3].

KARMEN 2 collected data from February 1997 through March 2001 corresponding to 9425 C accumulated proton charge on the ISIS target. A spatial coincidence between the initial  $e^+$  and the neutron capture of  $1.3 \text{m}^3$  was required. Applying all cuts to the data [10], 15  $(e^+, n)$  candidate sequences were finally reduced. Figure 3 shows the remaining sequences in the appropriate energy, time and spatial windows. The background components are also given with their distributions. All components except the intrinsic  $\bar{\nu}_e$  contamination are measured online in different time and energy windows (see Table II). The extracted number of sequences is in excellent agreement with the background expectation, consistent with no oscillation signal. To also include the detailed spectral information of each individual event, a maximum likelihood method is applied.

Analogously to the LSND likelihood function, we can define the combined likelihood for

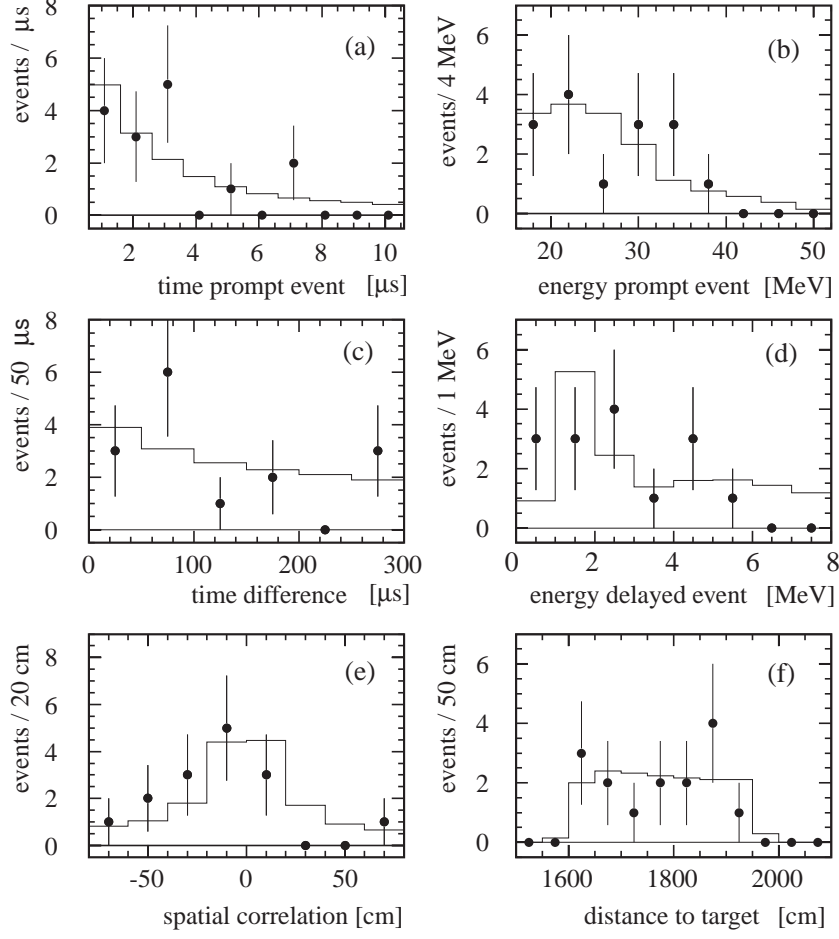


FIG. 3: KARMEN 2 event ensemble: (a) time of prompt events, (b) energy of prompt events, (c) time difference between prompt and delayed event, (d) energy of delayed events, (e) spatial correlation and (f) distance to target of prompt event. The 15 oscillation candidates are in very good agreement with the background expectation of 15.8 events (solid line).

the KARMEN sample as

$$\mathcal{L} = \left( \prod_{j=1}^{N_{beam-on}} p_j \right), \quad (2.8)$$

where

$$p_j(\vec{x}_j) = \sum_{i=1}^{N_{contr.}} q_i(\vec{x}_j) \cdot r_i \quad (2.9)$$

with  $N_{beam-on} = 15$  and  $N_{contr.} = 5$ , the oscillation signal as well as 4 background contributions. For KARMEN 2, the vector  $\vec{x}$  has, of course, a different definition, namely  $\vec{x} = (E_{pr}, T_{pr}, E_{del}, \Delta T, \Delta x)$  describing the energy and time relative to beam-on-target of the prompt event, the energy of the delayed event as well as the time difference and spatial

TABLE II: Expected KARMEN 2 background components and  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  signal

Process	Expectation	Determination
Cosmic induced background	$3.9 \pm 0.2$	measured in diff. time window
Charged current coincidences	$5.1 \pm 0.2$	measured in diff. energy, time windows
$\nu_e$ -induced random coincidences	$4.8 \pm 0.3$	measured in diff. time window
$\bar{\nu}_e$ source contamination	$2.0 \pm 0.2$	MC-simulation
Total background $N_{bg}$	$15.8 \pm 0.5$	
$N_{osc}(\sin^2(2\Theta) = 1, \Delta m^2 = 100 \text{ eV}^2/c^4)$	$2913 \pm 267$	

correlation of the prompt and delayed event. Again, normalizations (2.3) and (2.4) with the appropriate event parameters have to hold. Differing from the likelihood definition in the LSND analysis, for KARMEN with its very small background, all background components are combined into one contribution, therefore reducing the above sum (2.9) to only the oscillation signal and the total background,  $N_{contr.} = 2$ . Furthermore, the weight for the background variation, defined in (2.5) as a Gaussian factor, is taken as Gamma function, reflecting a continuous variation in a Poisson statistics [10].

Figure 4 shows the logarithm of the likelihood as a function of the oscillation parameters  $(\sin^2(2\Theta), \Delta m^2)$ . The maximum is reached near the physical boundary  $\sin^2(2\Theta) = 0$ , i.e. the no-oscillation case. The sharp drop in  $\ln \mathcal{L}$  towards larger values of  $\sin^2(2\Theta)$  demonstrates that there is no indication for an oscillation signal.

A typical approach to deduce confidence regions from likelihood functions would consist of defining the contour of an area of given confidence by cutting the likelihood function at the corresponding value below the maximum. For a Gaussian shaped likelihood function of two uncorrelated free parameters, the corresponding values would be

$$\Delta \ln \mathcal{L}_{\text{Gauss}} = -\ln(1 - \delta/100) = 2.3(4.6) \quad (2.10)$$

for  $\delta = 90(99)\%$  confidence, respectively [11]. However, this method leads to correct coverage only in the Gaussian case. As can be seen from Figures 2 and 4 the likelihood functions have distinct features: An oscillatory behavior as a function of the free parameters, the maximum at the physical boundary (for KARMEN), or numerous side maxima with comparable likelihood values (for LSND). In addition, the parameter space in  $(\sin^2(2\Theta), \Delta m^2)$  is,

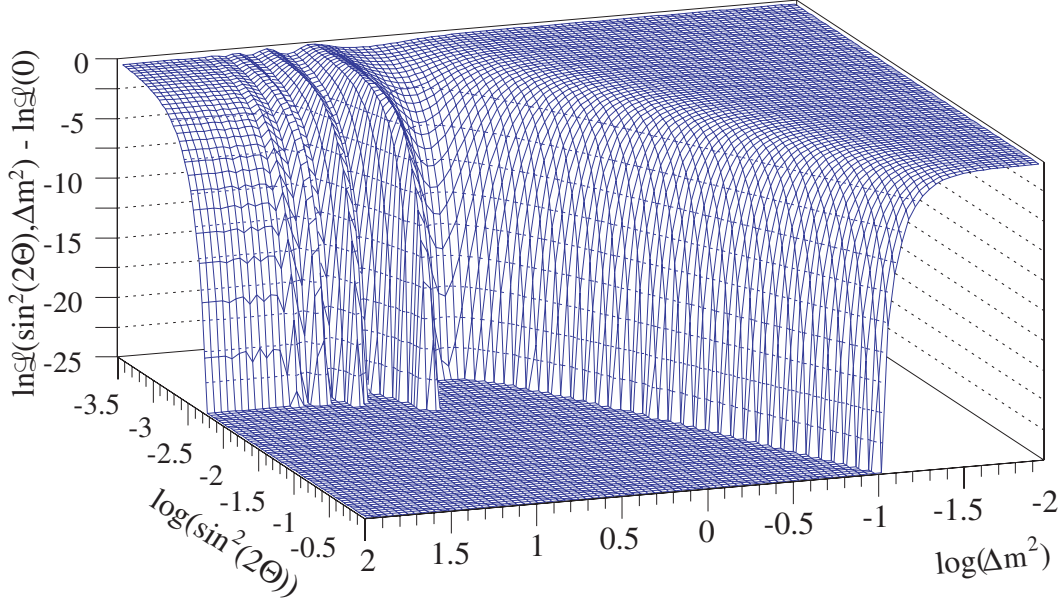


FIG. 4: KARMEN 2 logarithmic likelihood function  $\ln \mathcal{L}(\sin^2(2\Theta), \Delta m^2)$

in principle, not limited, its metric in terms of prior probability density not unequivocally defined.

### III. INDIVIDUAL CONFIDENCE REGIONS

The method we apply in the following is based on a unified frequentist approach suggested by [4] which eliminates the bias that occurs when one decides, after analysing the data, between using a confidence interval (having a positive signal) or an upper confidence limit (having a result compatible with the background). When first presented, it was argued that the suggested ordering principle near the physical boundary should be modified [12],[13], or that a Bayesian approach to the extracted likelihood function would be more appropriate [14]. However, many experimental results have been analysed since using the unified approach [15], also being described as a standard procedure by [16].

The basic idea of the unified frequentist approach is to create a large number of event samples in perfect analogy to an experiment. These samples are created by Monte Carlo using the full event information for the likelihood procedure. For an oscillation hypothesis  $H$  with given parameters  $(\sin^2(2\Theta), \Delta m^2)_H$ , a large sample of Monte Carlo simulations of so-called toy experiments is created. These simulations are based on the detailed knowl-

edge of all experimental resolution functions and the spectral information on the individual background contributions. In addition, they comprise the expected experimental signal for the given oscillation hypothesis [31]. For each likelihood function of the analysed samples, the estimator  $\Delta \ln \mathcal{L} = \ln \mathcal{L}_{max} - \ln \mathcal{L}_H$  is calculated by comparing each sample's likelihood maximum with the value for the given input parameters  $(\sin^2(2\Theta), \Delta m^2)_H$ . The hypothesis  $H$  is then accepted at a confidence  $\delta$ , if the estimator  $\Delta \ln \mathcal{L}$  of the experimental likelihood function is contained within the smallest  $\delta\%$  of the simulated estimator distribution. Since, in principle, the estimator distribution itself is a function of the oscillation parameters, a complete statistical analysis consists of a scan of the entire parameter space  $(\sin^2(2\Theta), \Delta m^2)$  to extract the according region of confidence.

### A. KARMEN

To deduce confidence regions from the KARMEN 2 likelihood function, the parameter space  $(\sin^2(2\Theta), \Delta m^2)$  has been scanned extracting the estimator distribution for 8000 different oscillation hypotheses, including the no-oscillation scenario. From these distributions, each containing the information of about 4000 simulated and analysed samples, regions of different confidence levels can be calculated. Figure 5 shows the regions of the oscillation parameters  $(\sin^2(2\Theta), \Delta m^2)$  for confidence levels of 68 %, 90 % and 95 %. Parameter combinations to the right contain the complementary confidence. The contours of the confidence areas are often referred to as exclusion curves of a given confidence. For values  $\Delta m^2 \geq 100 \text{ eV}^2/c^4$ , the contours cut at oscillation amplitudes of  $\sin^2(2\Theta) = 0.9 \cdot 10^{-3}$  (68 % C.L.),  $\sin^2(2\Theta) = 1.7 \cdot 10^{-3}$  (90% C.L.) and  $\sin^2(2\Theta) = 2.2 \cdot 10^{-3}$  (95% C.L.).

The KARMEN 2 sensitivity, defined as expectation value for the upper limit of a given confidence interval under the assumption of no oscillations, is again obtained by simulations:  $\langle \sin^2(2\Theta) \rangle = 0.6(1.6, 2.2) \cdot 10^{-3}$  for 68 % C.L. (90% C.L., 95% C.L.), respectively. The error on the contour lines is given by the statistical error of the estimator's cut value  $\Delta \ln \mathcal{L}(\delta)$  and therefore depends on the amount of simulated samples per hypothesis and the degree  $\delta$  of confidence under consideration. For example, determining  $\Delta \ln \mathcal{L}(90\%)$  from a distribution of 1000 simulated samples relies on the upper 10 % tail or the highest 100  $\Delta \ln \mathcal{L}$  values. We derive the error on  $\Delta \ln \mathcal{L}$  from the spread of the cut value taking the upper  $100 \pm \sqrt{100}$  samples. For KARMEN 2, the typical relative error on  $\Delta \ln \mathcal{L}(90\%)$  is in the range of 2–3 %.

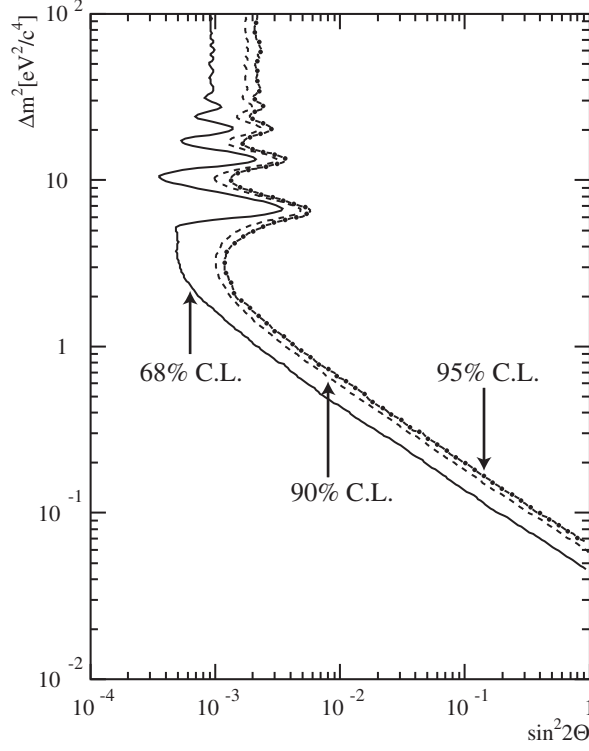


FIG. 5: KARMEN 2 regions of various confidence. Parameter combinations to the right contain the complementary confidence.

## B. LSND

For the LSND likelihood function, the method to deduce the confidence regions is identical to the one described in section III A for KARMEN 2. Technically, however, there are differences due to the large amount of computing time. This arises from the fact that the event samples are much larger (about 1000 events, depending on the oscillation parameters, instead of about 15) and the integration over different background contributions according to equation (2.5) is performed in as much as 255 discrete steps of different backgrounds BUB and BRB. Therefore, we restricted the hypotheses tested from a fine grid for KARMEN 2 to some grid points representing the typical variation of expected oscillation events. For a fixed oscillation parameter  $\Delta m^2 = 1 \text{ eV}^2/\text{c}^4$ , 4 oscillation scenarios were simulated with 1000 individual samples each (see Table III).

As verified by earlier calculations [5] and the KARMEN 2 analysis, the main variation of  $\Delta \ln \mathcal{L}(\delta)$  for a given confidence  $\delta$  in the parameter space  $(\sin^2(2\Theta), \Delta m^2)$  follows the

TABLE III: LSND estimator values  $\Delta \ln \mathcal{L}$  for 3 different confidence levels  $\delta$  and varying input oscillation signal strength  $\sin^2(2\Theta)$  but fixed parameter  $\Delta m^2 = 1 \text{ eV}^2/\text{c}^4$ .

$\sin^2(2\Theta)$	$N_{osc}$	$\Delta \ln \mathcal{L}$		
	DAR+DIF	68 %	90 %	95 %
0.0	0.0	$2.25 \pm 0.2$	$4.25^{+0.3}_{-0.5}$	$5.45^{+1.2}_{-1.0}$
$2.5 \cdot 10^{-3}$	$33.5 \pm 0.2$	$2.55 \pm 0.1$	$4.05^{+0.7}_{-0.4}$	$5.15^{+1.8}_{-0.7}$
$5 \cdot 10^{-3}$	$67.1 \pm 0.4$	$2.65 \pm 0.1$	$4.25^{+0.5}_{-0.4}$	$5.35^{+1.1}_{-0.8}$
$1 \cdot 10^{-2}$	$134.2 \pm 0.8$	$1.95 \pm 0.1$	$3.55^{+0.7}_{-0.4}$	$4.55^{+1.9}_{-0.9}$

variation of the signal strength, i.e. the number of oscillation events  $N_{osc}$  among the event sample. We therefore extrapolated the values of  $\Delta \ln \mathcal{L}$  in Table III linearly with  $N_{osc}$  and adopted, at each grid point  $(\sin^2(2\Theta), \Delta m^2)$  with its expected number of oscillation events  $N_{osc}$ , the estimator value  $\Delta \ln \mathcal{L}(\sin^2(2\Theta), \Delta m^2)$ . Having constructed the estimator cut distribution this way, the last step is analogous to that used for KARMEN2, i.e. to cut the LSND likelihood function with the given values  $\Delta \ln \mathcal{L}(\sin^2(2\Theta), \Delta m^2)$  below the maximum and draw the contour lines. Figure 6 shows the contour lines of areas containing 68 %, 90 % and 95 % confidence. Although parameter combinations with  $\Delta m^2 > 10 \text{ eV}^2/\text{c}^4$  are part of the 90% C.L. area, they are scarcely included in the more stringent interval of 68 % confidence, demonstrating the lower likelihood values for such oscillation scenarios.

It is useful to compare the extracted confidence regions with the favored regions given in [3] where the likelihood function had been calculated based on a larger event sample (5697 events with  $20 < E_e < 200 \text{ MeV}$  and no  $R_\gamma$ -cut). In addition, a simplified extraction of confidence regions according to equation (2.10) had been applied [32]. Figure 7 shows, that for lower values of  $\Delta m^2$ , there is only a slight shift to larger oscillation amplitudes of the new analysis. A large overlap of the 90% C.L. area with the region according to  $\Delta \ln \mathcal{L} = 2.3$  demonstrates the good statistical agreement of the analyses of both event samples. For higher values of  $\Delta m^2$ , the differences are more pronounced. Compared to the older analysis, the confidence region is not only shifted considerably to larger amplitudes, but the likelihood values themselves are larger relative to the likelihood maximum which results in a larger area of 90% C.L.. These effects can be understood by the upper energy cut  $E_e < 60 \text{ MeV}$  applied in the actual analysis. Restricting the analysis mainly to the oscillation channel

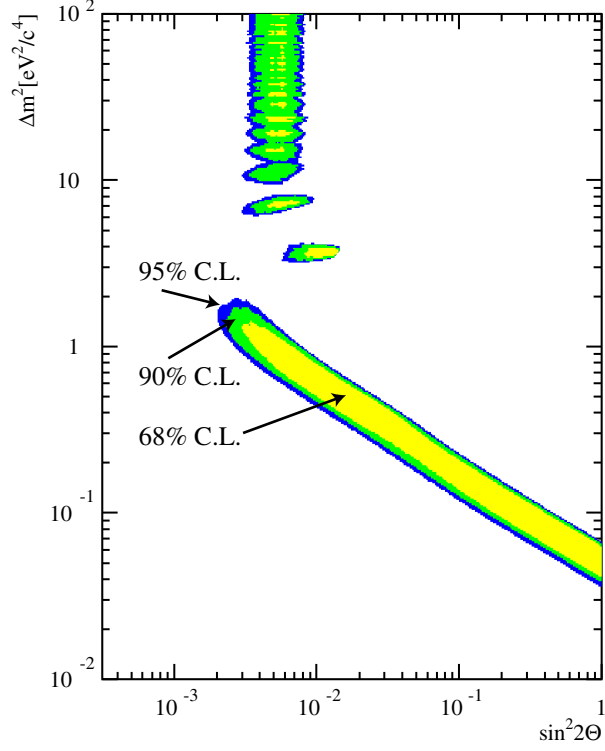


FIG. 6: LSND regions of various confidence. Parameter combinations outside the contours contain the complementary confidence.

$\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  favors also large values of  $\Delta m^2$  whereas a signal from  $\nu_\mu \rightarrow \nu_e$ , derived in [3] to be much smaller, but detectable mainly at higher energies, decreases the likelihood for such solutions as well as the absolute strength or oscillation amplitude.

Finally, we note the importance to derive values of  $\Delta \ln \mathcal{L}$  leading to correct coverage of confidence regions. These values differ considerably from the simplified Gaussian approach of a constant of 2.3 units. In an earlier analysis of the LSND data [5], the typical estimator cut values were  $\Delta \ln \mathcal{L}(90\%) \approx 2.5 - 3.5$  on the basis of about 3000 events. Now, with even smaller event samples, the values of  $\Delta \ln \mathcal{L}$  further increase to  $\approx 3.5 - 4.5$ . This dependance of  $\Delta \ln \mathcal{L}$  is also underlined by the typical values for the much smaller KARMEN 2 sample size, where  $\Delta \ln \mathcal{L}(90\%)$  ranges from 3 to 5.



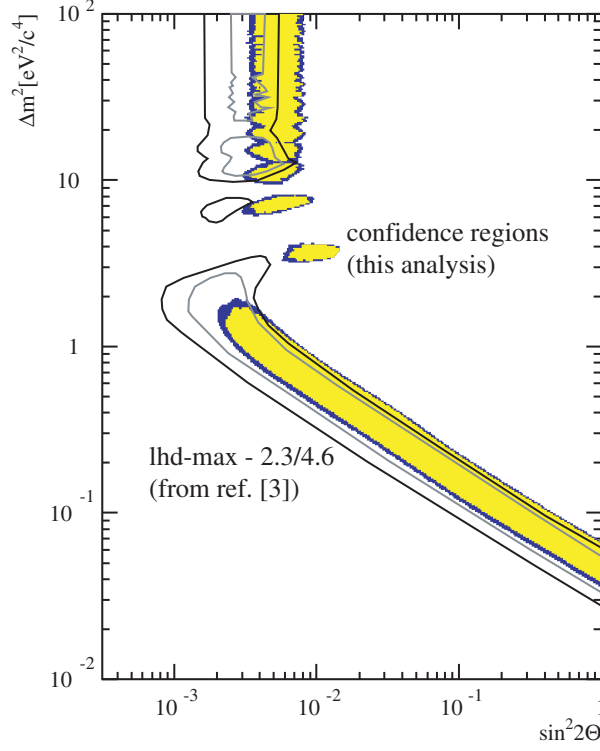


FIG. 7: LSND regions of 90 % and 95 % confidence in comparison with the favored region analysing the likelihood function of a larger event sample and applying a constant cut  $\Delta \ln \mathcal{L} = 2.3(4.6)$ , from [3]

#### IV. COMPATIBILITY ANALYSIS

Having analysed both experiments' likelihood functions with the same consistent method, we can now use this method and its results to deduce quantitative statements on the question of statistical compatibility of both experimental results and, in the case of such compatibility, on the common parameter combinations  $(\sin^2(2\Theta), \Delta m^2)$ .

##### A. Level of compatibility

One of the most misleading but nevertheless very frequently used interpretation of the LSND and KARMEN results is to take the LSND region left over from the KARMEN exclusion curve as area of  $(\sin^2(2\Theta), \Delta m^2)$  favored by both experiments. Such an interpretation, though appealingly straight forward, completely ignores the information of both likelihood

functions and reduces them to two discrete levels of a specific confidence  $\delta$ . To be able to correctly combine the two experimental results and extract the combined confidence regions, we have to use the original estimator distributions for KARMEN and for LSND. In a first step, the level of statistical compatibility of the two experimental outcomes has to be quantified. In the second step, the parameters favored by both experiments will be deduced.

In the following analysis, an *experiment* or *event*, in logical terms of frequency or probability of occurrence, means repeating both, the KARMEN *and* the LSND experiment. Compatibility is then achieved if there are at least some oscillation parameters being the element of both the repeated KARMEN-like and LSND-like experiment's confidence region.

We make the well justified assumption that the two experiments LSND and KARMEN are independent. Then, a two dimensional estimator distribution can be constructed for each hypothesis  $(\sin^2(2\Theta), \Delta m^2)_H$  from the individual distributions  $\Delta \ln \mathcal{L}_K$  of KARMEN and  $\Delta \ln \mathcal{L}_L$  of LSND by an inverse projection. For each hypothesis, the pair of experimental values  $(\Delta \ln \mathcal{L}_{KARMEN2}, \Delta \ln \mathcal{L}_{LSND})$  is checked to be part of the 'inner'  $\delta\%$  of the simulated 2-dimensional estimator distribution. There are different methods of ordering these 2-dimensional distributions all leading to very similar results [5]. In the following, we require each experimental value  $\Delta \ln \mathcal{L}_{KARMEN2}$  and  $\Delta \ln \mathcal{L}_{LSND}$  to be contained in the corresponding 1-dimensional interval of confidence. This selection criterion is equivalent to selecting a rectangle of frequency of the 2-dim estimator distribution. The resulting confidence is  $\delta = \delta_{KARMEN} \cdot \delta_{LSND}$ . Since we do not weight the experiments, we define  $\delta_{KARMEN} = \delta_{LSND}$ , so that for a combined confidence of e.g. 81 %, we demand  $\Delta \ln \mathcal{L}_{KARMEN2} < \Delta \ln \mathcal{L}_K(90\%)$  and  $\Delta \ln \mathcal{L}_{LSND} < \Delta \ln \mathcal{L}_L(90\%)$ .

At a level of  $\delta_{KARMEN} = \delta_{LSND} = 60\%$ , no parameter combination  $(\sin^2(2\Theta), \Delta m^2)$  fulfills the requirement  $(\Delta \ln \mathcal{L}_{KARMEN2} < \Delta \ln \mathcal{L}_K) \wedge (\Delta \ln \mathcal{L}_{LSND} < \Delta \ln \mathcal{L}_L)$  any more. This corresponds to the fact that, at the level of combined confidence

$$\delta_{inc} = 0.6^2 = 36\% , \quad (4.1)$$

the two experiments are completely incompatible. Coming back to our definition of probability of occurrence, in 64 % of repetitions of double experiments, a KARMEN2 plus an LSND experiment with their typical parameters and statistics, the outcome would consist of statistically compatible results with a specified set of oscillation parameters  $(\sin^2(2\Theta), \Delta m^2)$ .

## B. Common oscillation parameters

In the following, we analyse the preferred regions of oscillation parameters in the case of compatibility defined in the section before. If two experiments are independent producing two likelihood functions as result of the event analysis, the recipe for a combined analysis is straight forward. The likelihoods can be multiplied, i.e. in our case the logarithms of the likelihood functions are added. The absolute values of the likelihood functions are somewhat arbitrary. A recommended presentation of  $\ln \mathcal{L}$  is to normalize the individual functions  $\ln \mathcal{L}_{KARMEN2}$  and  $\ln \mathcal{L}_{LSND}$  to a point in  $(\sin^2(2\Theta), \Delta m^2)$  where they are equally sensitive to a potential signal [17]. In our case of the oscillation search this corresponds to values of  $\sin^2(2\Theta) = 0$ . Therefore we define the combined likelihood function

$$\begin{aligned} \ln \mathcal{L}(\sin^2(2\Theta), \Delta m^2) = & \left\{ \ln \mathcal{L}_{KARMEN2}(\sin^2(2\Theta), \Delta m^2) - \ln \mathcal{L}_{KARMEN2}(\sin^2(2\Theta) = 0) \right\} \\ & + \left\{ \ln \mathcal{L}_{LSND}(\sin^2(2\Theta), \Delta m^2) - \ln \mathcal{L}_{LSND}(\sin^2(2\Theta) = 0) \right\} . \end{aligned} \quad (4.2)$$

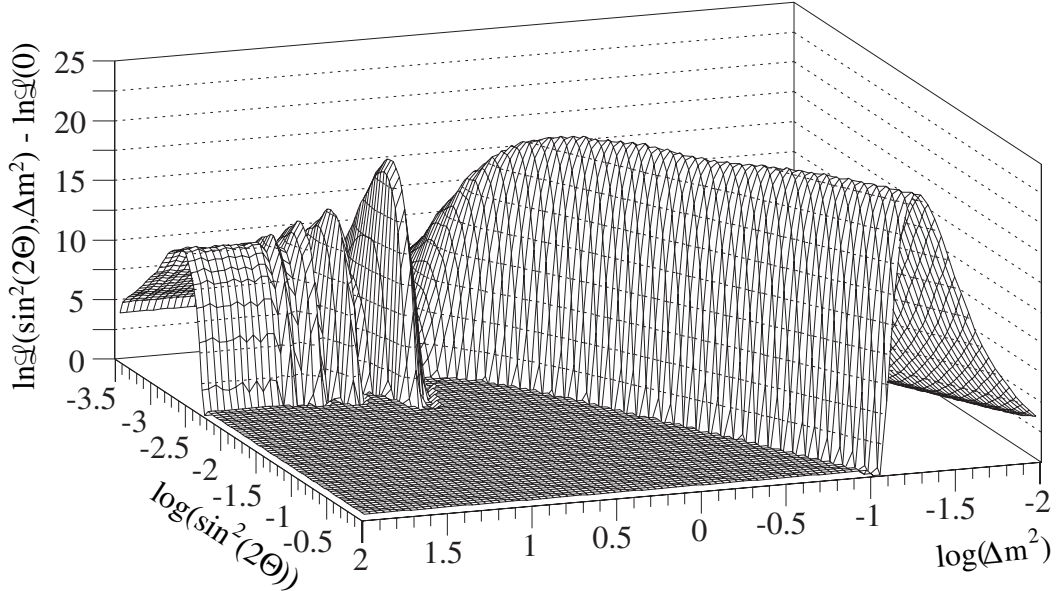


FIG. 8: Combined KARMEN 2 and LSND logarithmic likelihood function  $\ln \mathcal{L}(\sin^2(2\Theta), \Delta m^2)$

Figure 8 shows the combined function  $\ln \mathcal{L}(\sin^2(2\Theta), \Delta m^2)$  with a maximum of  $\ln \mathcal{L}(\sin^2(2\Theta)=1, \Delta m^2=0.05)=21.5$  on a long flat 'ridge' of low  $\Delta m^2$  values. The positive signal of LSND dominates the combined likelihood. However, the maximum in comparison

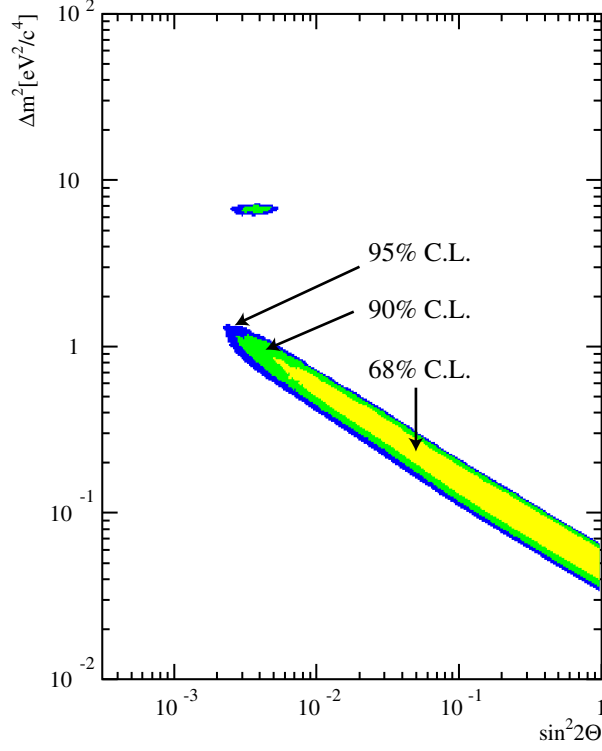


FIG. 9: Regions of various confidence for the combined analysis assuming statistical compatibility of KARMEN 2 and LSND.

to the no-oscillation value is reduced (see Equ. 2.7) as well as the shape at higher values of  $\Delta m^2$  is significantly modified due to the KARMEN 2 likelihood function.

To determine the confidence regions in  $(\sin^2(2\Theta), \Delta m^2)$ , again we apply the unified approach, i.e. we create the estimator distribution  $\Delta \ln \mathcal{L}_{K+L}$  for various hypotheses  $(\sin^2(2\Theta), \Delta m^2)_H$  and compare the experiment's value  $\Delta \ln \mathcal{L}_{KARMEN2+LSND}$  with the cut value  $\Delta \ln \mathcal{L}_{K+L}(\delta)$  for a given confidence. To get the estimator, a simulated event sample now consists of a KARMEN 2-like and an LSND-like MC sample each analysed with the appropriate likelihood definitions. Then both likelihood functions are added following equation (4.2) and the estimator  $\Delta \ln \mathcal{L}_{K+L} = \ln \mathcal{L}_{K+L}(\sin^2(2\Theta), \Delta m^2)_{max} - \ln \mathcal{L}_{K+L}(\sin^2(2\Theta), \Delta m^2)_H$  is calculated. The typical cut values for a confidence of  $\delta = 90\%$  are slightly higher than the individual LSND ones, ranging from 3.5 to 5.5. Figure 9 shows the confidence regions of the oscillation parameters for the combined likelihood analysis. The total confidence of a parameter region is hereby reduced by the fraction of incompatibility

of the two experiments, so

$$\delta_{tot} = \delta \cdot (1 - \delta_{inc}) \quad (4.3)$$

which results in e.g.  $\delta_{tot} = 0.9 \cdot 0.64 = 58\%$  total confidence for the parameter combinations within the area denoted as 90% C.L. area in Figure 9.

The combination of both experiments' results with this consistent frequentist approach demonstrates that  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  solutions with  $\Delta m^2 > 10 \text{ eV}^2/\text{c}^4$  are excluded. There remain essentially two solutions in the parameter space of  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  oscillations, one at  $\Delta m^2 \approx 7 \text{ eV}^2/\text{c}^4$  and the area with  $\Delta m^2 < 1 \text{ eV}^2/\text{c}^4$ . The latter one, though graphically rather large in Figure 9 corresponds to just one solution in the 'phase space' of oscillations: Any  $\bar{\nu}_e$  signal from oscillation has rather low energy, the evolution of the oscillation probability (equ. 1.1) being just at the beginning of the oscillation length. With an oscillation parameter  $\Delta m^2 = 0.5 \text{ eV}^2/\text{c}^4$ , and a typical lower energy of  $E_\nu = 20 \text{ MeV}$  from  $\mu^+$  decay at rest, the first maximum of oscillations of  $\bar{\nu}_\mu$  into  $\bar{\nu}_e$  would be at a distance

$$d_{1.max} = \frac{\pi}{2} \cdot \frac{E_\nu}{1.27 \cdot \Delta m^2} \approx 50 \text{ m} \quad (4.4)$$

from the neutrino source. For such oscillation parameters, the negative result of KARMEN 2 compared to the excess of LSND mainly reflects the different detector distances  $d_{KARMEN} \approx 17 \text{ m}$  versus  $d_{LSND} \approx 30 \text{ m}$  from the source. This is one way to reconcile the results of the two experiments.

To summarize, at a joint confidence level of  $\delta_{tot} = 0.36 + 0.64 \cdot 0.68 \approx 80\%$ , the two experiments are either incompatible or lead to just one single oscillation solution at  $\Delta m^2 < 1 \text{ eV}^2/\text{c}^4$ . The remaining 20% confidence separate into an enlargement of the parameter region of low  $\Delta m^2$  values and into a second oscillation solution at  $\Delta m^2 \approx 7 \text{ eV}^2/\text{c}^4$ .

## V. CONCLUSION

We have analysed both the KARMEN 2 and the LSND final data with a maximum likelihood method using a similar event-based likelihood function. The likelihood functions as functions of the free parameters  $\sin^2(2\Theta)$  and  $\Delta m^2$  demonstrate the different outcomes of the experiments, namely a clear, statistically significant excess of  $\bar{\nu}_e$  in LSND versus no indication of an oscillation signal in KARMEN 2. To deduce regions of correct coverage, we

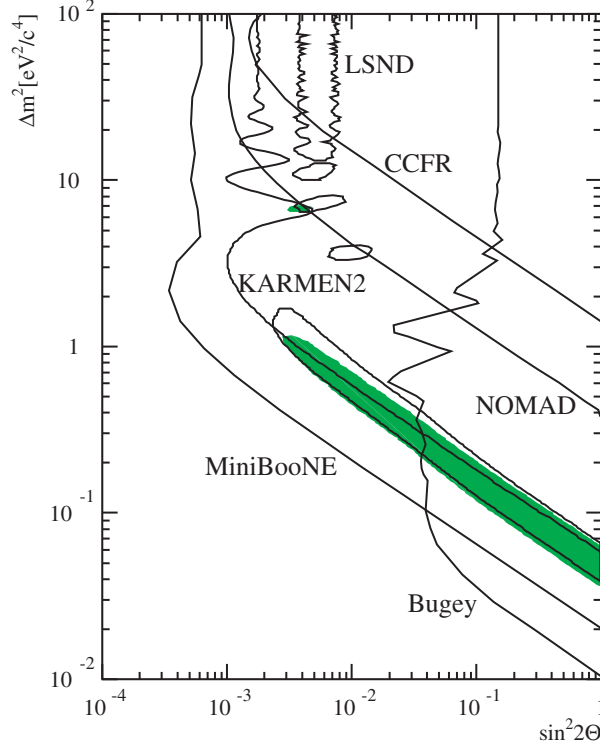


FIG. 10: Parameter regions deduced in this work (grey area) compared with existing limits of experiments (Bugey  $\bar{\nu}_e \rightarrow \bar{\nu}_x$  [18], CCFR  $\nu_\mu \rightarrow \nu_e$  [19] and NOMAD  $\nu_\mu \rightarrow \nu_e$  [20]) and the envisaged sensitivity of the MiniBooNE experiment (with final single horn design [21]).

applied a unified frequentist approach to both likelihood analyses individually. The results underline the feasibility of as well as the necessity for such an approach.

A quantitative joint statistical analysis has been performed leading to a level of 36 % incompatibility of the experimental outcomes, corresponding to individual confidence levels of 60 %. For the cases of statistical compatibility, the common parameter regions have been identified on the basis of the unified frequentist approach applied to the combined likelihood function of KARMEN 2 and LSND. The derived confidence regions in  $(\sin^2(2\Theta), \Delta m^2)$  clearly differ from an often applied but incorrect graphical overlap of the confidence regions of the individual experiments. There are two oscillation scenarios with either  $\Delta m^2 \approx 7 \text{ eV}^2/\text{c}^4$  or  $\Delta m^2 < 1 \text{ eV}^2/\text{c}^4$  compatible with both experiments.

We performed a joint statistical analysis incorporating some of the systematic uncertainties of the experiments, such as neutrino flux uncertainty, accuracy of known cross sections and resolution functions of both experiments. Further –unknown– systematic uncertainties

might become evident only by performing new experiments to confirm or discard the results described here. Figure 10 shows the intended sensitivity of a new experiment, MiniBooNE at Fermilab [21], which is under construction and will independently crosscheck the LSND evidence.

In addition, we did not incorporate the results of other experiments on the oscillation parameters investigated. As shown in Figure 10, the deduced favored region in  $(\sin^2(2\Theta), \Delta m^2)$  is partly overlapped by 90% C.L. exclusion curves of other experiments. In particular, the recent NOMAD  $\nu_\mu \rightarrow \nu_e$  result [20] clearly reduces the overall likelihood for the  $\Delta m^2 \approx 7 \text{ eV}^2/c^4$  solution. A complete analysis should include these results on the basis of the same statistical method, i.e. a consistent frequentist analysis. This implies, however, the detailed knowledge of experimental data and resolution functions of these experiments not accessible to us. Furthermore, the exclusion curve from the Bugey experiment is based on the disappearance search  $\bar{\nu}_e \rightarrow \bar{\nu}_x$ . Combining this experiment correctly with the appearance results of  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  or  $\nu_\mu \rightarrow \nu_e$  in terms of mixing angles would therefore also require a full three- or four-dimensional mixing scheme with theoretical models including sterile neutrinos [22, 23, 24, 25, 26, 27] or CPT violation [28] and would become model-dependent.

## VI. ACKNOWLEDGEMENTS

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- [1] S. Fukuda *et al.* (Super-Kamiokande), Phys. Rev. Lett. **85**, 3999 (2000);  
M. Ambrosio *et al.* (MACRO), Phys. Lett B **517**, 59 (2001).
  - [2] B. T. Cleveland *et al.* (Homestake), Astrophys. J. **496**, 505 (1998);  
W. Hampel *et al.* (Gallex), Phys. Lett. B **447**, 127 (1999);  
J. N. Abdurashitov *et al.* (SAGE), Phys. Rev. C **60**, 055801 (1999);  
S. Fukuda *et al.* (Super-Kamiokande), Phys. Rev. Lett. **86**, 5651 (2001); **86**, 5656 (2001);  
M. Altmann *et al.* (GNO), Phys. Lett. B **490**, 16 (2000);  
Q. R. Achmad *et al.* (SNO), Phys. Rev. Lett. **87**, 071301 (2001).
  - [3] A. Aguilar *et al.*, Phys. Rev. D **64**, 112007 (2001).
  - [4] G. J. Feldman and R. D. Cousins, Phys. Rev. D **57**, 3873 (1998).
  - [5] K. Eitel, New Jour. Phys. **2**, 1.1 (2000).
  - [6] C. Athanassopoulos *et al.*, Phys. Rev. C **54**, 2685 (1996).
  - [7] R. L. Burman, M. E. Potter, and E. S. Smith, Nucl. Instrum. Methods A **291**, 621 (1990);  
R. L. Burman, A. C. Dodd, and P. Plischke, Nucl. Instrum. Methods A **368**, 416 (1996).
  - [8] G. Drexlin *et al.*, Nucl. Instrum. Methods A **289**, 490 (1990);  
G. Drexlin, Progr. Part. Nucl. Phys **40**, 193 (1998).
  - [9] C. Athanassopoulos *et al.*, Nucl. Instrum. Methods A **388**, 149 (1997).
  - [10] B. Armbruster *et al.*, Phys. Rev. D (submitted), *Los Alamos e-print xxx.lanl.gov* hep-ex/0203021.
  - [11] For reviews, see L. Lyons, *Statistics for nuclear and particle physicists* (Cambridge University Press, 1986);  
W. T. Eadie *et al.*, *Statistical Methods in Experimental Physics* (North Holland, Amsterdam and London, 1971).
  - [12] C. Giunti, Phys. Rev. D **59**, 053001 (1999).
  - [13] B. P. Roe, M. B. Woodroffe, Phys. Rev. D **60**, 053009 (1999).
  - [14] G. D'Agostini, *Bayesian Reasoning in High-Energy Physics and Applications* (CERN Report CERN-99-03).
  - [15] F. James *et al.*, *Proceedings of the Workshop on Confidence Limits, CERN, Geneva, 2000* (CERN Report CERN-2000-005).



- [16] D. E. Groom *et al.*, Eur. Phys. J. C **15**, 1 (2000).
- [17] G. D'Agostini, *Los Alamos e-print xxx.lanl.gov* hep-ex/0002055.
- [18] B. Achkar *et al.*, Nucl. Phys. B **434**, 503 (1995).
- [19] A. Romosan *et al.*, Phys. Rev. Lett. **78**, 2912 (1997).
- [20] V. Valuev, *Proceedings of the International Europhysics Conference on HEP* 2001.
- [21] A. O. Bazarko, Nucl. Phys. B (Proc. Suppl.) **91**, 210 (2001).
- [22] S. M. Bilenky *et al.*, Progr. Part. Nucl. Phys **43**, 1 (1999).
- [23] M. Hirsch and J. W. F. Valle, Phys. Lett. B **495**, 121 (2000).
- [24] A. Ioannisian and J. W. F. Valle, Phys. Rev. D **63**, 073002 (2001).
- [25] E. Ma and G. Rajasekaran, Phys. Rev. D **64**, 117303 (2001).
- [26] A. Perez-Lorenzana and C. A. de S. Pires, Phys. Lett. B **522**, 297 (2001).
- [27] A. Strumia, *Los Alamos e-print xxx.lanl.gov* hep-ph/0201134.
- [28] G. Barenboim *et al.*, Phys. Rev. D **65**, 053001 (2002).
- [29] In a comparison of different oscillation searches, one generally has to use a complete 3- or 4-dimensional description of neutrino masses and mixings, leading to the general oscillation probability. Since both experiments under investigation here have similar parameters and, most important, search for oscillations in the same appearance mode, the simplified formula holds for the purpose of the following analysis.
- [30] The spatial coordinate  $z$  along the detector axis can be easily referred to the neutrino flight path  $L$  by adding the distance source-center of detector tank.
- [31] For each individual sample, the number of background events as well as the number of signal events vary according to the error of the expectation values.
- [32] In fact, due in part to the larger statistical sample, in [3] the results of such a simplified procedure were shown to be similar to the results obtained from an approximation to a full unified frequentist approach.